

# Some Curious Consequences of the Minimal Length Uncertainty Relation

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## Work in collaboration with:

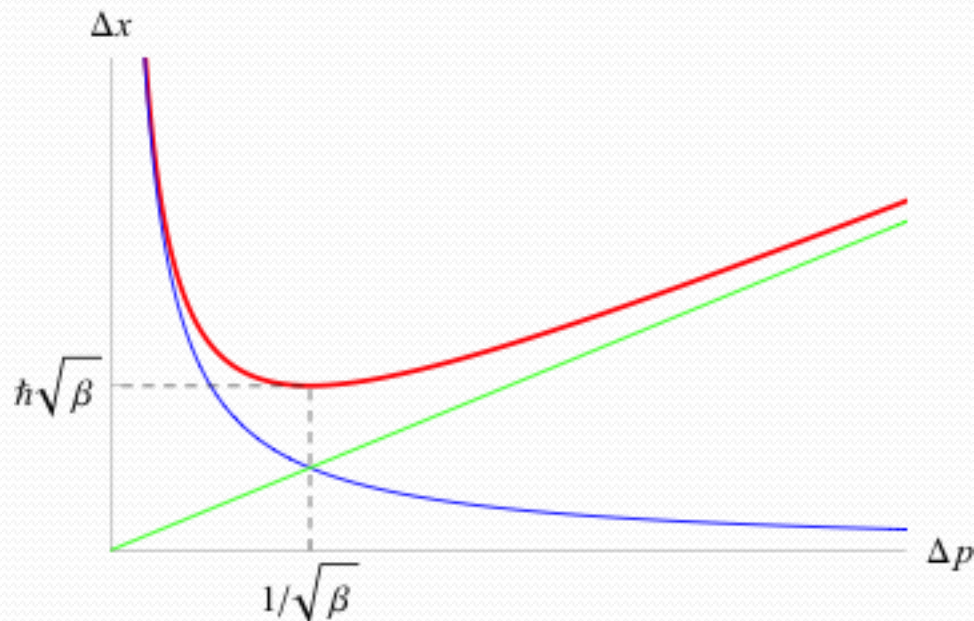
- Zack Lewis
- Djordje Minic
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## Also contributions from:

- Naotoshi Okamura (Yamanashi Univ.)
- Sandor Benczik (quit physics)
- Saif Rayyan (MIT, physics education)

# The Minimal Length Uncertainty Relation

$$\Delta x \geq \frac{\hbar}{2} \left( \frac{1}{\Delta p} + \beta \Delta p \right) \Rightarrow \Delta x \geq \Delta x_{\min} = \hbar \sqrt{\beta}$$



Suggested by Quantum Gravity. Observed in perturbative String Theory.

# Deformed Commutation Relation

$$\frac{1}{i\hbar} [\hat{x}, \hat{p}] = 1 + \beta \hat{p}^2 \quad \Rightarrow \quad \Delta x \geq \frac{\hbar}{2} \left( \frac{1}{\Delta p} + \beta \Delta p \right)$$

The operators can be represented as:

$$\begin{cases} \hat{x} = i\hbar(1 + \beta \hat{p}^2) \frac{d}{dp} \\ \hat{p} = p \end{cases}$$

and the inner product as

$$\langle f | g \rangle = \int_{-\infty}^{\infty} \frac{dp}{(1 + \beta p^2)} f^*(p) g(p)$$

# Harmonic Oscillator

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{1}{2}k\hat{x}^2$$
$$\Rightarrow \left[ -\frac{\hbar^2 k}{2} \left\{ (1 + \beta p^2) \frac{d}{dp} \right\}^2 + \frac{p^2}{2m} \right] \psi(p) = E \psi(p)$$

Can be solved exactly. Energy eigenvalues :

$$E_n = \frac{k}{2} \left[ \left( n + \frac{1}{2} \right) \sqrt{(\Delta x_{\min})^4 + 4a^4} + \left( n^2 + n + \frac{1}{2} \right) (\Delta x_{\min})^2 \right]$$

$$a = \sqrt[4]{\frac{\hbar^2}{km}} = \sqrt{\frac{\hbar}{m\omega}}$$

No longer evenly spaced.  $n^2$  - dependence is introduced.

## Multi Dimensional Case

$$\frac{1}{i\hbar} [\hat{x}_i, \hat{p}_j] = (1 + \beta \hat{p}^2) \delta_{ij} + \gamma \hat{p}_i \hat{p}_j$$

$$[\hat{p}_i, \hat{p}_j] = 0$$

$$\frac{1}{i\hbar} [\hat{x}_i, \hat{x}_j] = -\left\{ (2\beta - \gamma) + \beta(2\beta + \gamma) \hat{p}^2 \right\} \hat{L}_{ij}, \quad \hat{L}_{ij} = \frac{\hat{x}_i \hat{p}_j - \hat{x}_j \hat{p}_i}{1 + \beta \hat{p}^2}$$

The operators can be represented as:

$$\begin{cases} \hat{x}_i = i\hbar \left[ (1 + \beta \hat{p}^2) \frac{\partial}{\partial p_i} + \gamma p_i p_j \frac{\partial}{\partial p_j} + \left\{ \beta + \gamma \left( \frac{D+1}{2} \right) - \delta(\beta + \gamma) \right\} p_i \right] \\ \hat{p}_i = p_i \end{cases}$$

and the inner product as

$$\langle f | g \rangle = \int_{-\infty}^{\infty} \frac{dp}{[1 + (\beta + \gamma)p^2]^\delta} f^*(p) g(p)$$

# Isotropic Harmonic Oscillator in D dimensions

$$\Psi_D(p_1, p_2, \dots, p_D) = R(p) Y_{l_{D-2} m_{D-3} \dots l_{m_2} m_1}(\Omega)$$

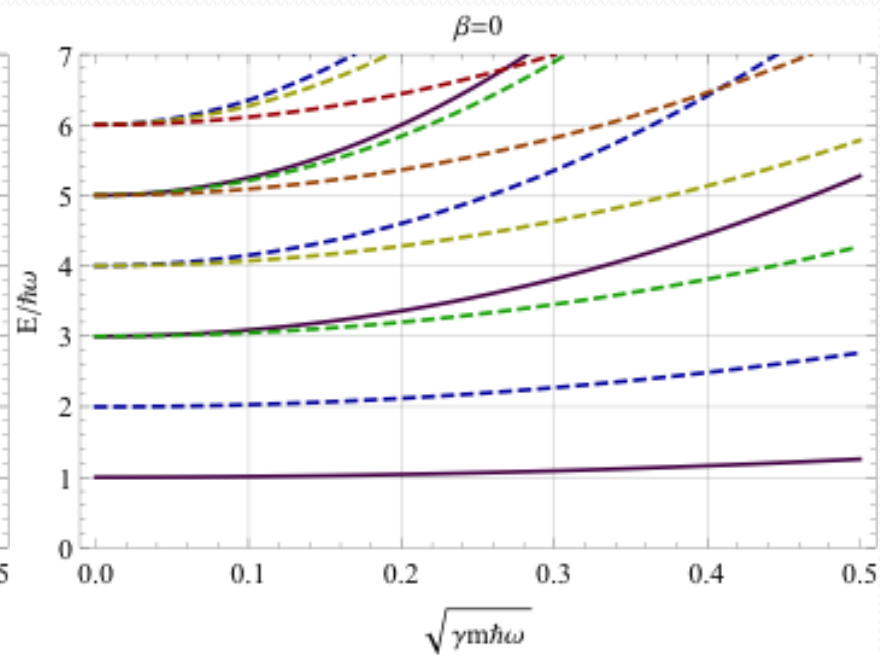
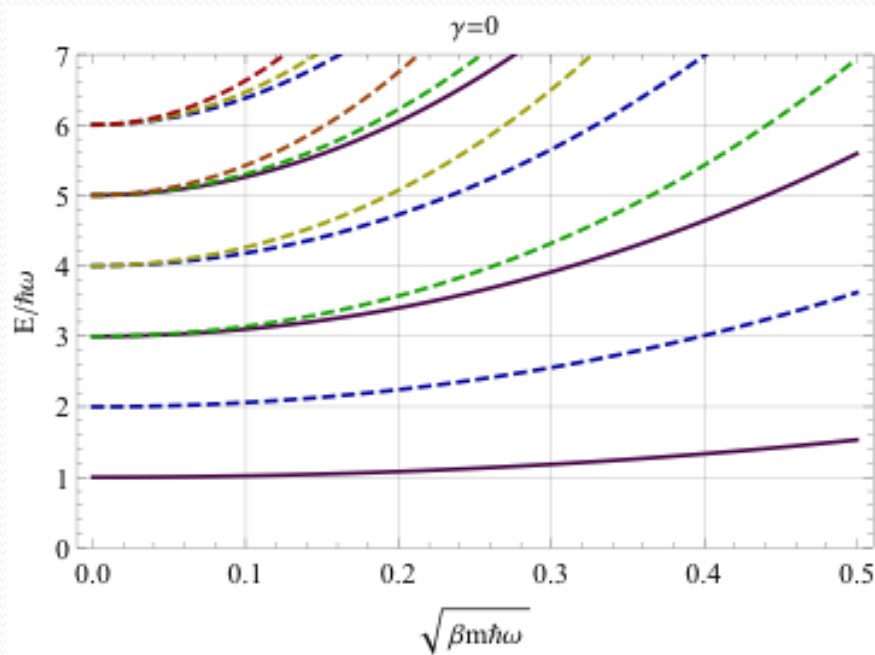
$$-\frac{\hbar^2 k}{2} \left[ \left\{ [1 + (\beta + \gamma) p^2] \frac{\partial}{\partial p} \right\}^2 + \frac{(D-1)(1 + \beta p^2)[1 + (\beta + \gamma) p^2]}{p} \frac{\partial}{\partial p} - \frac{L^2(1 + \beta p^2)^2}{p^2} \right] R(p) + \frac{p^2}{2m} R(p) = ER(p), \quad L^2 = l(l + D - 2)$$

Can be solved exactly. Energy eigenvalues :

$$E_{nl} = \frac{k}{2} \left[ \left( n + \frac{D}{2} \right) \sqrt{\Delta x_{\min}^4 [4L^2 + (D + \eta)^2]} + 4a^4 + \Delta x_{\min}^2 \left\{ (1 + \eta) \left( n + \frac{D}{2} \right)^2 + (1 - \eta) \left( L^2 + \frac{D^2}{4} \right) + \eta \frac{D}{2} \right\} \right], \quad \Delta x_{\min} = \hbar \sqrt{\beta}, \quad \eta = \frac{\gamma}{\beta}.$$

Dependence on angular momentum introduced. SU(D) degeneracy is broken.

# 2D Case





## Some Details (1D case) :

Change variable to :

$$\rho = \frac{1}{\sqrt{\beta}} \arctan \sqrt{\beta} p \quad -\frac{\pi}{2\sqrt{\beta}} < \rho < \frac{\pi}{2\sqrt{\beta}}$$

$$\left\{ \begin{array}{l} \hat{x} = i\hbar \frac{d}{d\rho} \\ \hat{p} = \frac{1}{\sqrt{\beta}} \tan \sqrt{\beta} \rho \end{array} \right. \quad \langle f | g \rangle = \int_{-\pi/2\sqrt{\beta}}^{\pi/2\sqrt{\beta}} d\rho f^*(\rho) g(\rho)$$

Schrodinger Equation :

$$\left[ -\frac{\hbar^2 k}{2} \frac{d^2}{d\rho^2} + \frac{1}{2m\beta} \tan^2 \sqrt{\beta} \rho \right] \psi(\rho) = E \psi(\rho)$$

Infinite square - well problem for large  $n$ , and also in the limit  $m \rightarrow \infty$ .

## Solution:

The solution is:

$$\psi_n^{(\lambda)}(\rho) = \sqrt[4]{\beta} \left[ 2^\lambda \Gamma(\lambda) \sqrt{\frac{n!(n+\lambda)}{2\pi\Gamma(n+2\lambda)}} \right] (\cos \sqrt{\beta}\rho)^\lambda C_n^\lambda(\sin \sqrt{\beta}\rho)$$

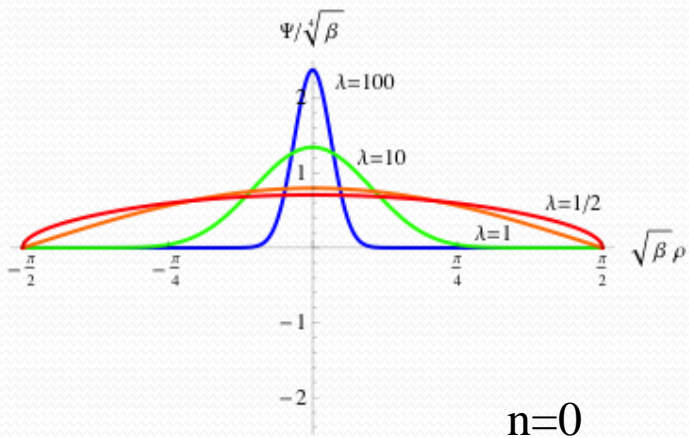
where:

$$\lambda = \frac{1}{2} + \sqrt{\frac{1}{4} + \left(\frac{a}{\Delta x_{\min}}\right)^4} \quad (1 < \lambda)$$

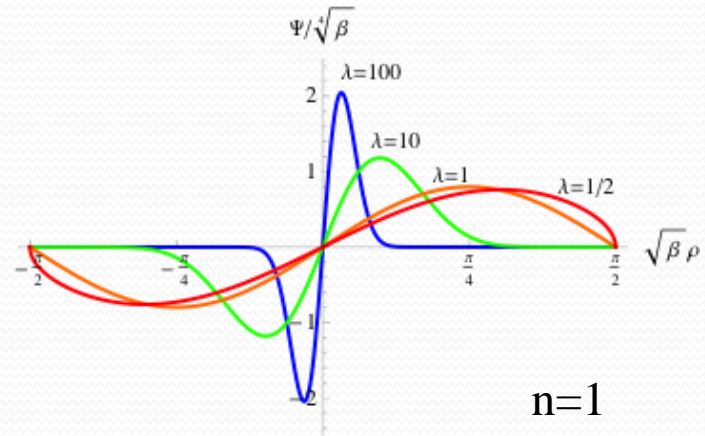
Uncertainties:

$$\Delta x = \Delta x_{\min} \sqrt{\frac{(\lambda+n)[(2\lambda-1)n+\lambda]}{(2\lambda-1)}}, \quad \Delta p = \frac{1}{\sqrt{\beta}} \sqrt{\frac{2n+1}{2\lambda-1}}$$

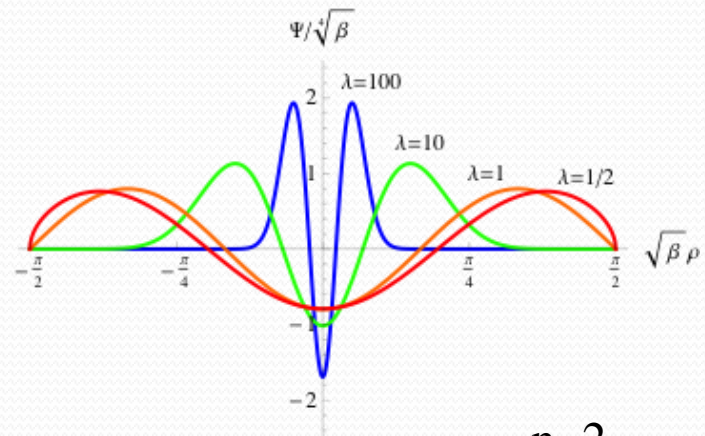
# Wave-functions:



$n=0$

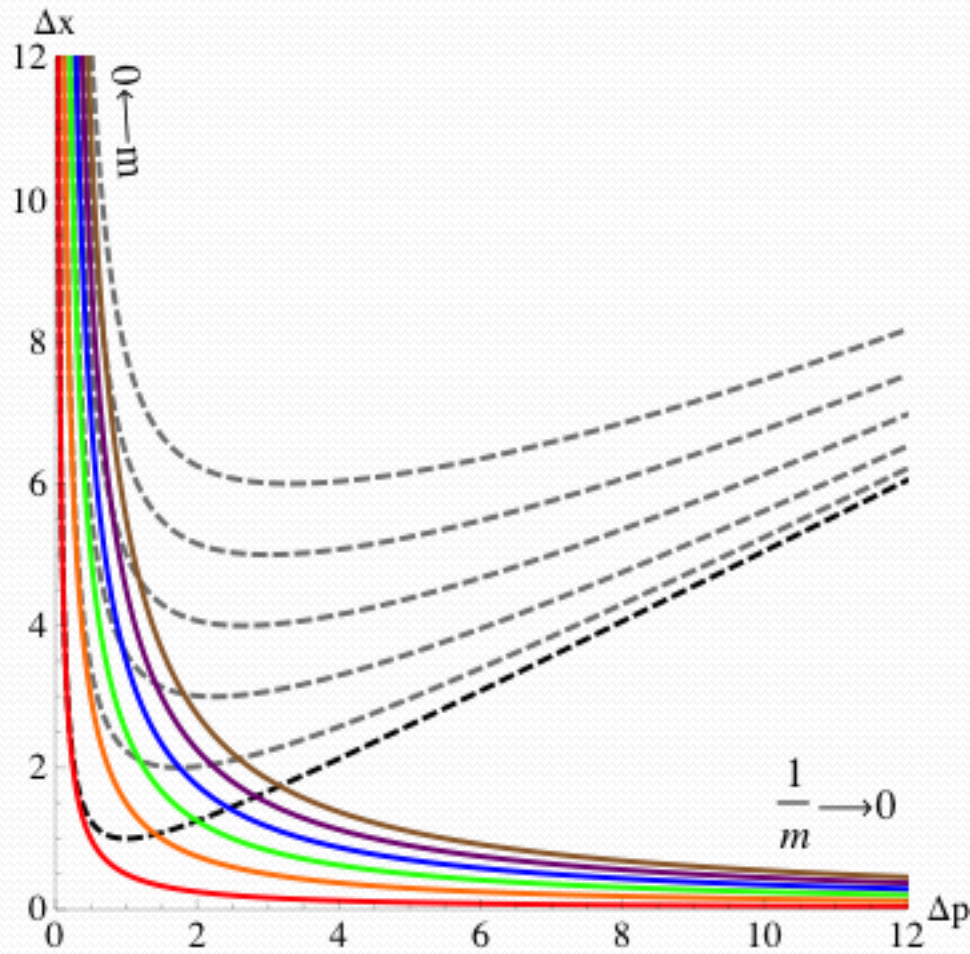


$n=1$

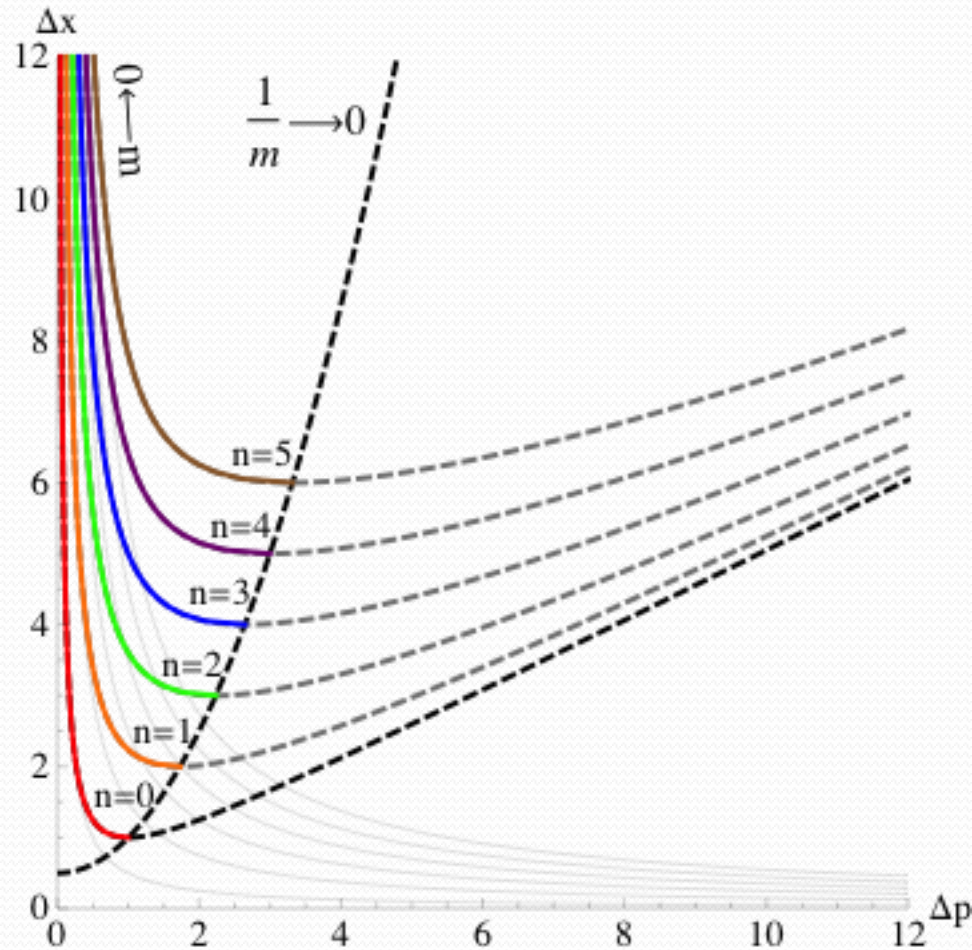


$n=2$

# Uncertainties of the Harmonic Oscillator: $\beta=0$



# Uncertainties of the Harmonic Oscillator: $\beta \neq 0$



How can we get onto the  $\Delta x \sim \Delta p$  branch?

## Harmonic Oscillator with negative mass:

$$\hat{H} = -\frac{\hat{p}^2}{2|m|} + \frac{1}{2}k\hat{x}^2$$
$$\Rightarrow \left[ -\frac{\hbar^2 k}{2} \left\{ (1 + \beta p^2) \frac{d}{dp} \right\}^2 - \frac{p^2}{2|m|} \right] \psi(p) = E \psi(p)$$

Energy eigenvalues :

$$E_n = \frac{k}{2} \left[ \left( n + \frac{1}{2} \right) \sqrt{(\Delta x_{\min})^4 - 4a^4} + \left( n^2 + n + \frac{1}{2} \right) (\Delta x_{\min})^2 \right]$$

$$a = \sqrt[4]{\frac{\hbar^2}{k|m|}}$$

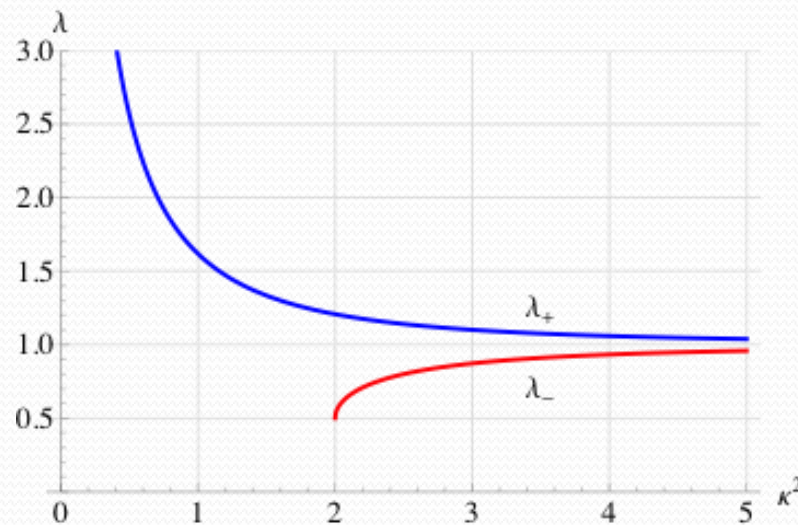
# Harmonic Oscillator with negative mass:

The solution is the same as before

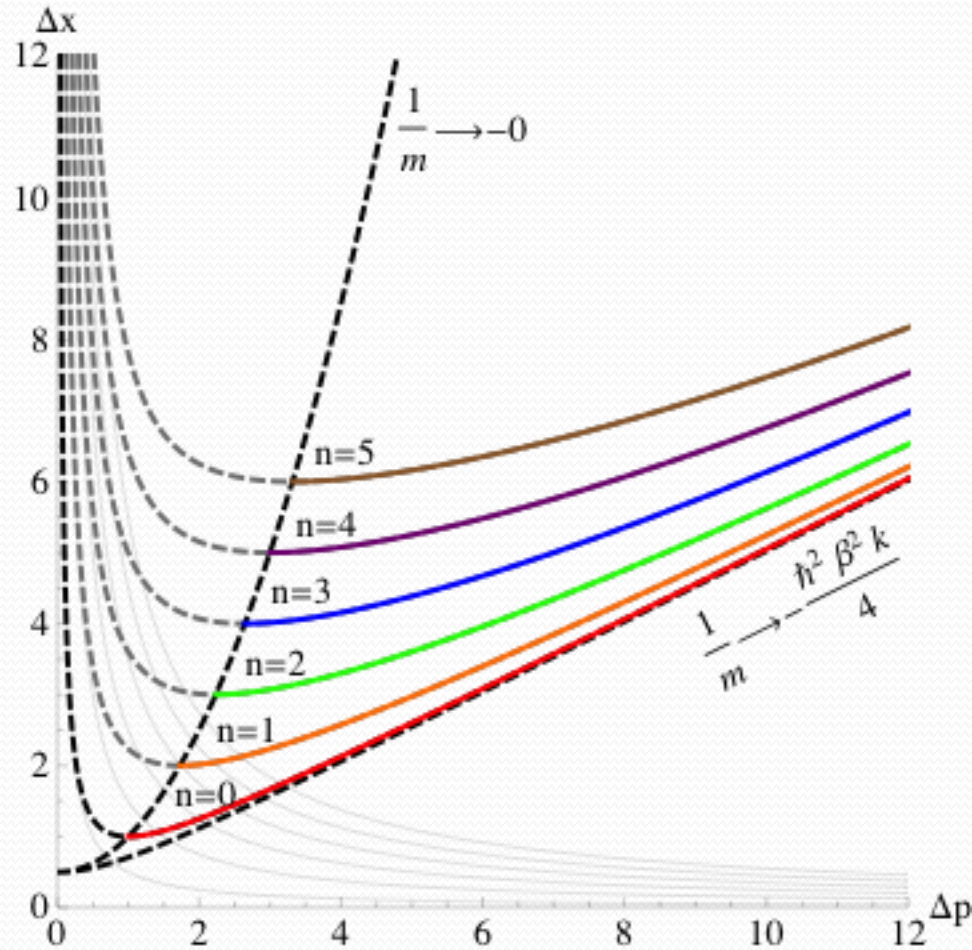
$$\psi_n^{(\lambda)}(\rho) = \sqrt[4]{\beta} \left[ 2^\lambda \Gamma(\lambda) \sqrt{\frac{n!(n+\lambda)}{2\pi\Gamma(n+2\lambda)}} \right] (\cos \sqrt{\beta}\rho)^\lambda C_n^\lambda(\sin \sqrt{\beta}\rho)$$

except :

$$\lambda = \frac{1}{2} + \sqrt{\frac{1}{4} - \left(\frac{a}{\Delta x_{\min}}\right)^4} \quad \left( \frac{1}{2} < \lambda < 1, \quad \Delta x_{\min} > \sqrt{2a} \right)$$



# Uncertainties of the Harmonic Oscillator: $\beta \neq 0$ , $m < 0$





## Classical Limit:

$$\frac{1}{i\hbar} [\hat{x}, \hat{p}] = (1 + \beta \hat{p}^2) \Rightarrow \{x, p\} = (1 + \beta p^2)$$

Classical Equations of Motion:

$$\dot{x} = \{x, H\}, \quad \dot{p} = \{p, H\}$$

Liouville Theorem :

$$dx \wedge dp \rightarrow \frac{dx \wedge dp}{1 + \beta p^2}$$

$h(1 + \beta p^2)$  can be considered a  $p$ -dependent effective  $h(p)$ .

# Classical Harmonic Oscillator:

Hamiltonian :

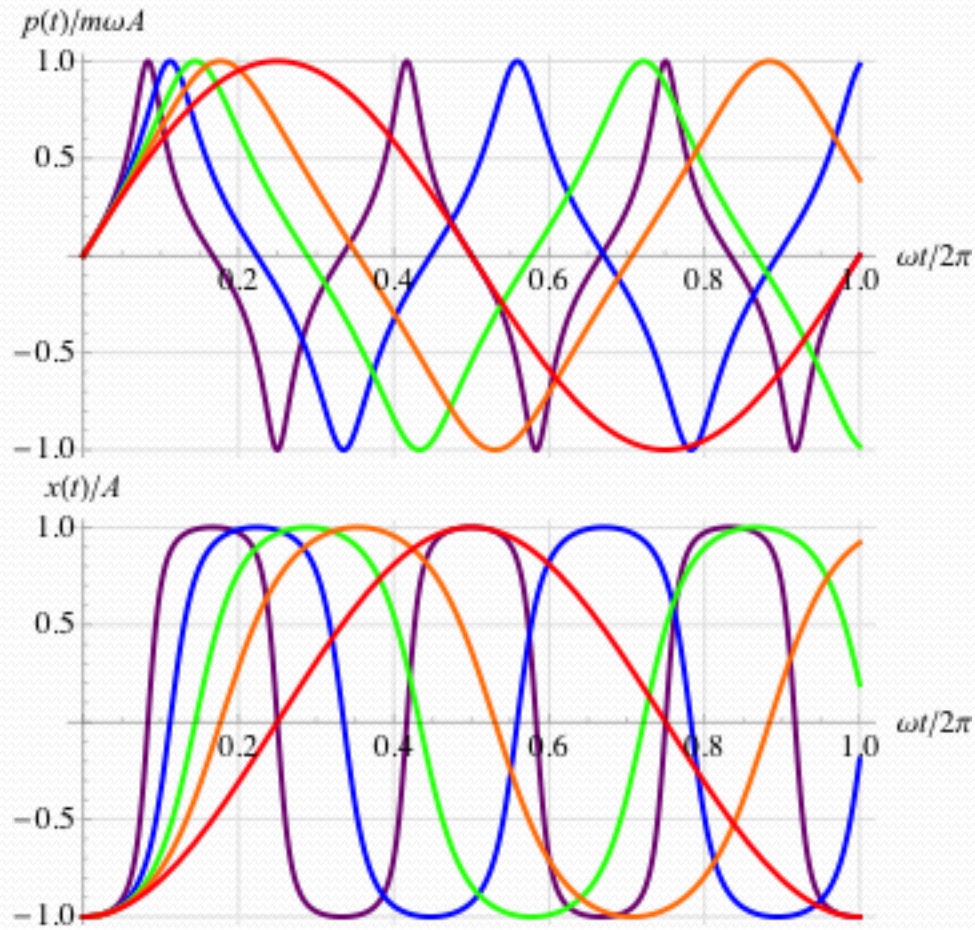
$$H = \frac{p^2}{2m} + \frac{1}{2} kx^2$$

Classical Equations of Motion:

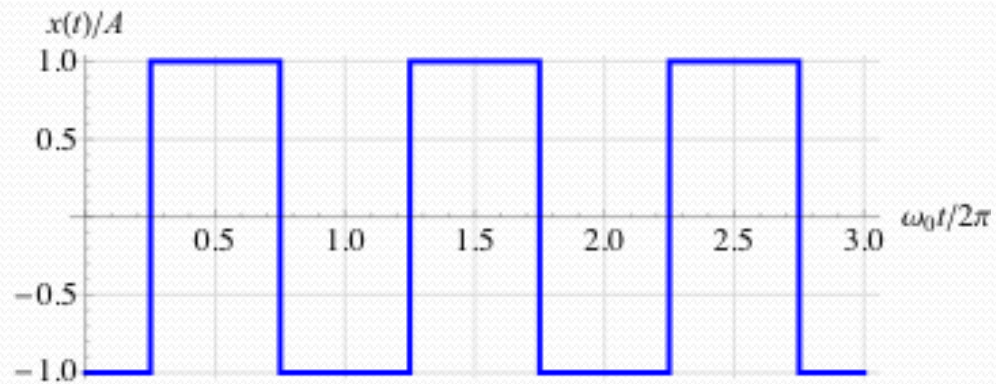
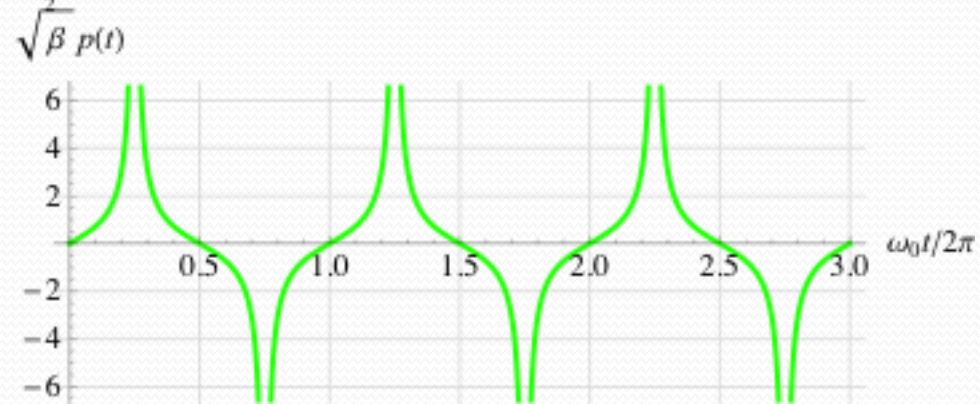
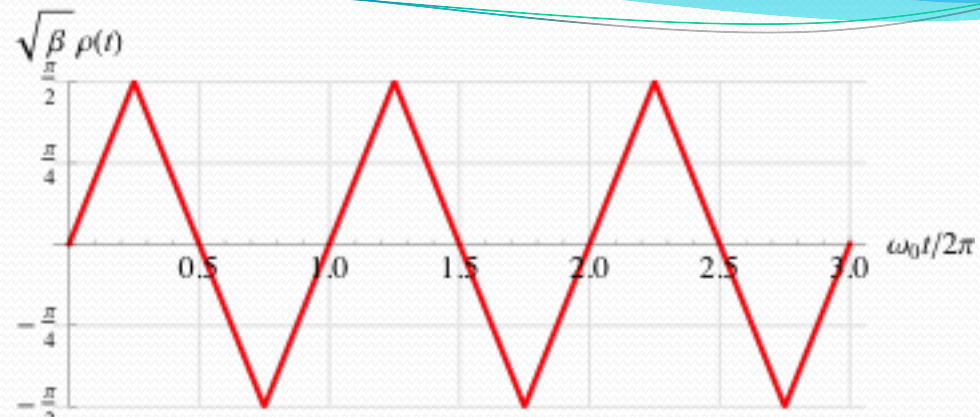
$$\begin{cases} \dot{x} = \{x, H\} = \frac{1}{m} (1 + \beta p^2) p \\ \dot{p} = \{p, H\} = -k(1 + \beta p^2) x \end{cases}$$

Time - dependence of  $x$  and  $p$  are different, but the trajectories in phase space are the same as the  $\beta = 0$  case since the Hamiltonian is the same.

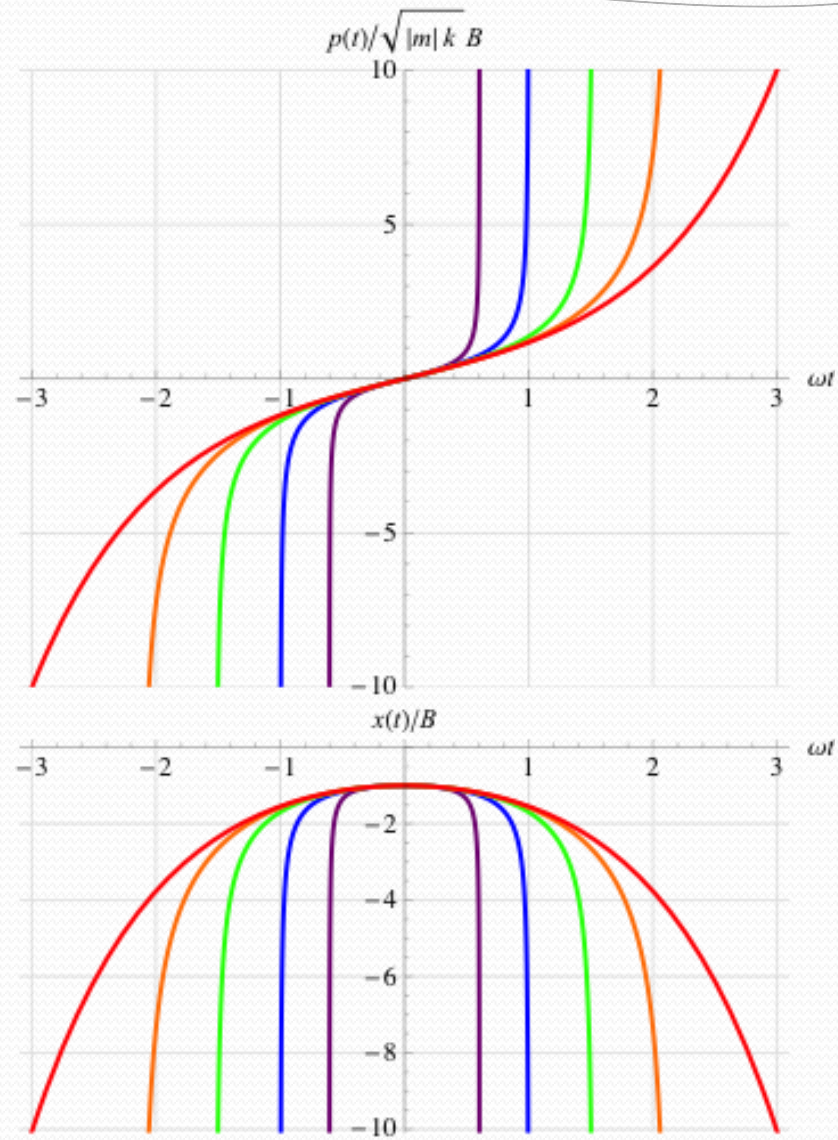
$m > 0$  case:



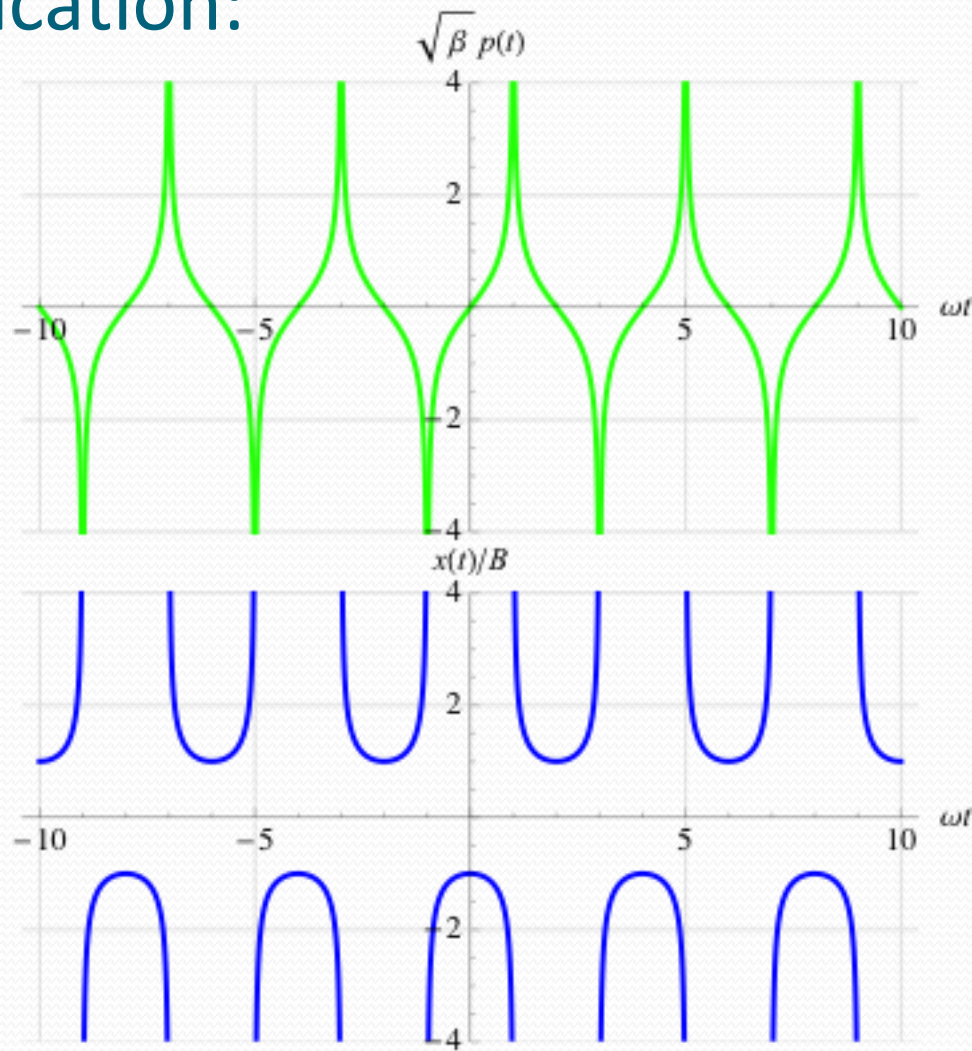
$m \rightarrow \infty$  limit:



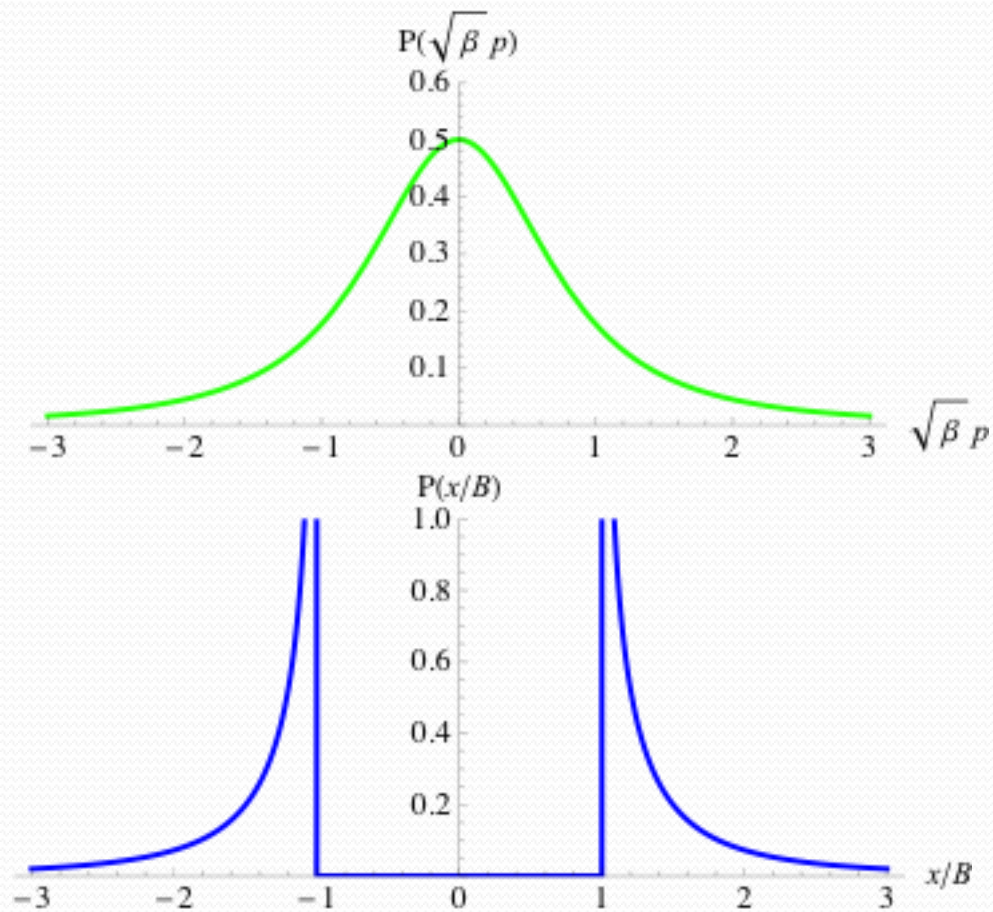
$m < 0$  case:



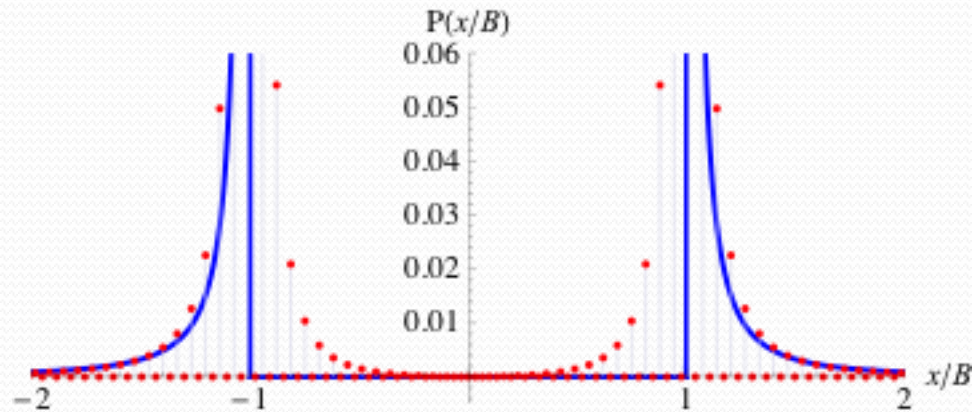
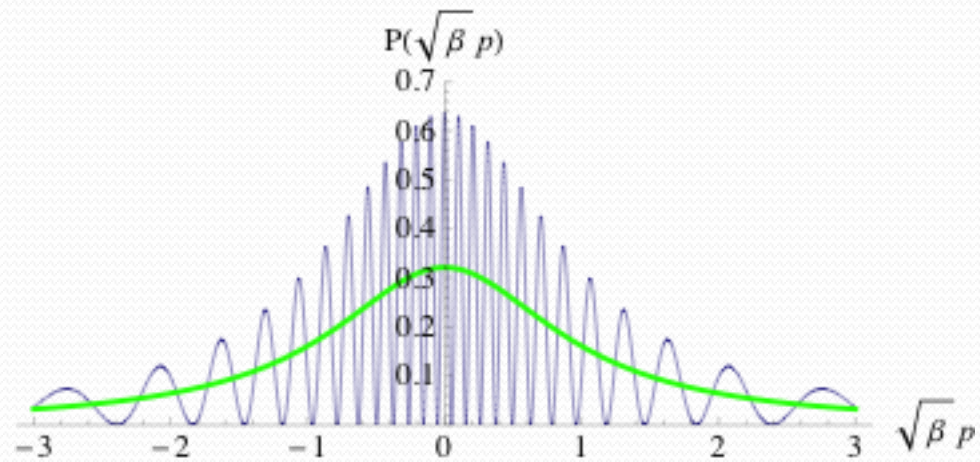
# Compactification:



# Classical Probabilities:



# Comparison with Quantum Probabilities:





## Work in progress:

Other potentials:

$$V(x) = F|x|$$

$$V(x) = \begin{cases} Fx & (x > 0) \\ \infty & (x < 0) \end{cases}$$

- Discrete energy eigenstates have been found for the negative mass case.
- Uncertainties are difficult to calculate. What do we mean by  $x=0$ ?

## Conclusions:

- The minimal length uncertainty relation allows discrete energy “bound” states for “inverted” potentials.
- In the classical limit, these “bound” states can be understood to be due to the finite time the particle spends near the phase-space origin.
- Particles move at arbitrary large velocities. Do the non-relativistic negative mass states correspond to relativistic imaginary mass tachyons?